PRESS RELEASE

The search for the missing gravitational signal

A new SISSA study proposes an array of interferometers in space to detect subtle fluctuations in the background gravitational signal that may reveal the secrets of black hole mergers

Trieste, 27 March 2023

Every year, hundreds of thousands of pairs of black holes merge in a cosmic dance that emits gravitational waves in every direction. Since 2015, the large ground-based LIGO, Virgo and KAGRA interferometers have made it possible to detect these signals, although only about a hundred such events, an infinitesimal fraction of the total, have been observed. Most of the waves remain "indistinguishable", superimposed and added together, creating a flat, diffuse background signal that scientists call the "stochastic gravitational wave background" (SGWB). New SISSA research, published in The Astrophysical Journal, proposes using a constellation of three or four space interferometers to map the flat and almost perfectly homogeneous background in a search for ripples. These small fluctuations, known to scientists as anisotropies, hold the information needed to understand the distribution of gravitational wave sources on the largest cosmological scale.
Researchers are convinced that next-generation detectors, such as the Einstein Telescope and the Laser Interferometer Space Antenna (LISA), will make direct measurement of the gravitational wave background possible in the foreseeable future. "Measuring these background fluctuations, known more correctly as anisotropies, will however continue to be extremely difficult, as identifying them requires a very high level of angular resolution not possessed by current and next generation survey instruments", explains Giulia Capurri, a SISSA PhD student and first author of the study.

Capurri, supervised by Carlo Baccigalupi and Andrea Lapi, has suggested that this problem could be overcome by means of a ‘constellation’ of three or four space interferometers in solar orbit and covering a distance approximating that between Earth and the Sun. With increasing separation, interferometers achieve better angular resolution, improving their ability to distinguish sources of gravitational waves. "A constellation of space interferometers orbiting the Sun could enable us to see subtle fluctuations in the gravitational background signal, thus allowing us to extract valuable information about the distribution of black holes, neutron stars and all other sources of gravitational waves in the universe" states Capurri.

Following the success of the LISA project’s space mission test, there are currently two proposals for the creation of space-based interferometer constellations: one European - the Big Bang Observatory (BBO), and one Japanese - the Deci-hertz Interferometer Gravitational-wave Observatory (DECIGO). “This represents one of the earliest work to provide specific predictions of the size of the stochastic background of gravitational waves by a constellation of instruments orbiting the Sun. Together with further similar projects whose details will be published in due course, they will be crucial for developing an optimal design for future observational instruments that we hope will be built and commissioned in the coming decades” concludes Carlo Baccigalupi, professor of theoretical cosmology at SISSA.

In the era of multimessenger astronomy, which began with ground-based interferometers such as LIGO and Virgo, the gravitational-wave background could pave the way to a new understanding of the universe on the large scale, as has already happened with the cosmic microwave background.
### USEFUL LINKS

**Full paper:** [The Astrophysical Journal](#)

### IMAGE

**Credits:** Simon Barke (University of Florida)

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